

N -Body Methods and Algorithms

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Basic Integration

Hermite Method

Neighbour Scheme

Close Encounters

Chain Regularization

Astrophysics

Post-Newtonian

Practicals

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Newton's Equation

$$\text{Force} \quad \mathbf{F}_i = -G \sum_{j=1; j \neq i}^N m_j \frac{\mathbf{r}_i - \mathbf{r}_j}{|\mathbf{r}_i - \mathbf{r}_j|^3}$$

Explicit differentiation

$$\begin{aligned} \dot{\mathbf{F}}_i = & -G \sum_{j=1; j \neq i}^N m_j \frac{\dot{\mathbf{r}}_i - \dot{\mathbf{r}}_j}{|\mathbf{r}_i - \mathbf{r}_j|^3} \\ & - 3m_j \frac{(\mathbf{r}_i - \mathbf{r}_j) \cdot (\dot{\mathbf{r}}_i - \dot{\mathbf{r}}_j)}{|\mathbf{r}_i - \mathbf{r}_j|^2} \frac{\mathbf{r}_i - \mathbf{r}_j}{|\mathbf{r}_i - \mathbf{r}_j|^3} \end{aligned}$$

New solution at $t = \Delta t$

$$\Delta \dot{\mathbf{r}}_i = \left(\frac{1}{2} \dot{\mathbf{F}}_i \Delta t + \mathbf{F}_i \right) \Delta t$$

$$\Delta \mathbf{r}_i = \left(\left(\frac{1}{6} \dot{\mathbf{F}}_i \Delta t + \frac{1}{2} \mathbf{F}_i \right) \Delta t + \dot{\mathbf{r}}_i \right) \Delta t$$

Repeat cycle for $i = 1, N$; N^2 interactions

Hermite Integration

Taylor series for \mathbf{F} and $\mathbf{F}^{(1)}$

$$\mathbf{F} = \mathbf{F}_0 + \mathbf{F}_0^{(1)} t + \frac{1}{2} \mathbf{F}_0^{(2)} t^2 + \frac{1}{6} \mathbf{F}_0^{(3)} t^3$$

$$\mathbf{F}^{(1)} = \mathbf{F}_0^{(1)} + \mathbf{F}_0^{(2)} t + \frac{1}{2} \mathbf{F}_0^{(3)} t^2$$

Prediction

$$\mathbf{r}_j = \left(\left(\frac{1}{6} \mathbf{F}_0^{(1)} \delta t'_j + \frac{1}{2} \mathbf{F}_0 \right) \delta t'_j + \mathbf{v}_0 \right) \delta t'_j + \mathbf{r}_0$$

$$\mathbf{v}_j = \left(\left(\frac{1}{2} \mathbf{F}_0^{(1)} \delta t'_j + \mathbf{F}_0 \right) \delta t'_j + \mathbf{v}_0 \right); \quad \delta t'_j = t - t_0$$

New forces $\mathbf{F}, \mathbf{F}^{(1)}$

Higher derivatives

$$\mathbf{F}_0^{(3)} = (2(\mathbf{F}_0 - \mathbf{F}) + (\mathbf{F}_0^{(1)} + \mathbf{F}^{(1)}) t) \frac{6}{t^3}$$

$$\mathbf{F}_0^{(2)} = (-3(\mathbf{F}_0 - \mathbf{F}) - (2\mathbf{F}_0^{(1)} + \mathbf{F}^{(1)}) t) \frac{2}{t^2}$$

Corrector for i

$$\Delta \mathbf{r}_i = \frac{1}{24} \mathbf{F}_0^{(2)} \Delta t^4 + \frac{1}{120} \mathbf{F}_0^{(3)} \Delta t^5$$

$$\Delta \mathbf{v}_i = \frac{1}{6} \mathbf{F}_0^{(2)} \Delta t^3 + \frac{1}{24} \mathbf{F}_0^{(3)} \Delta t^4$$

Basic Regularization

Two-body equation $\ddot{x} = -\frac{M}{x^2}$

Smoothing function $t' \equiv \frac{dt}{d\tau} = x$

Rule of differentiation $\frac{d}{dt} = \frac{1}{x} \frac{d}{d\tau}$

Time-smoothed equation $x'' = \frac{x'^2}{x} - M$

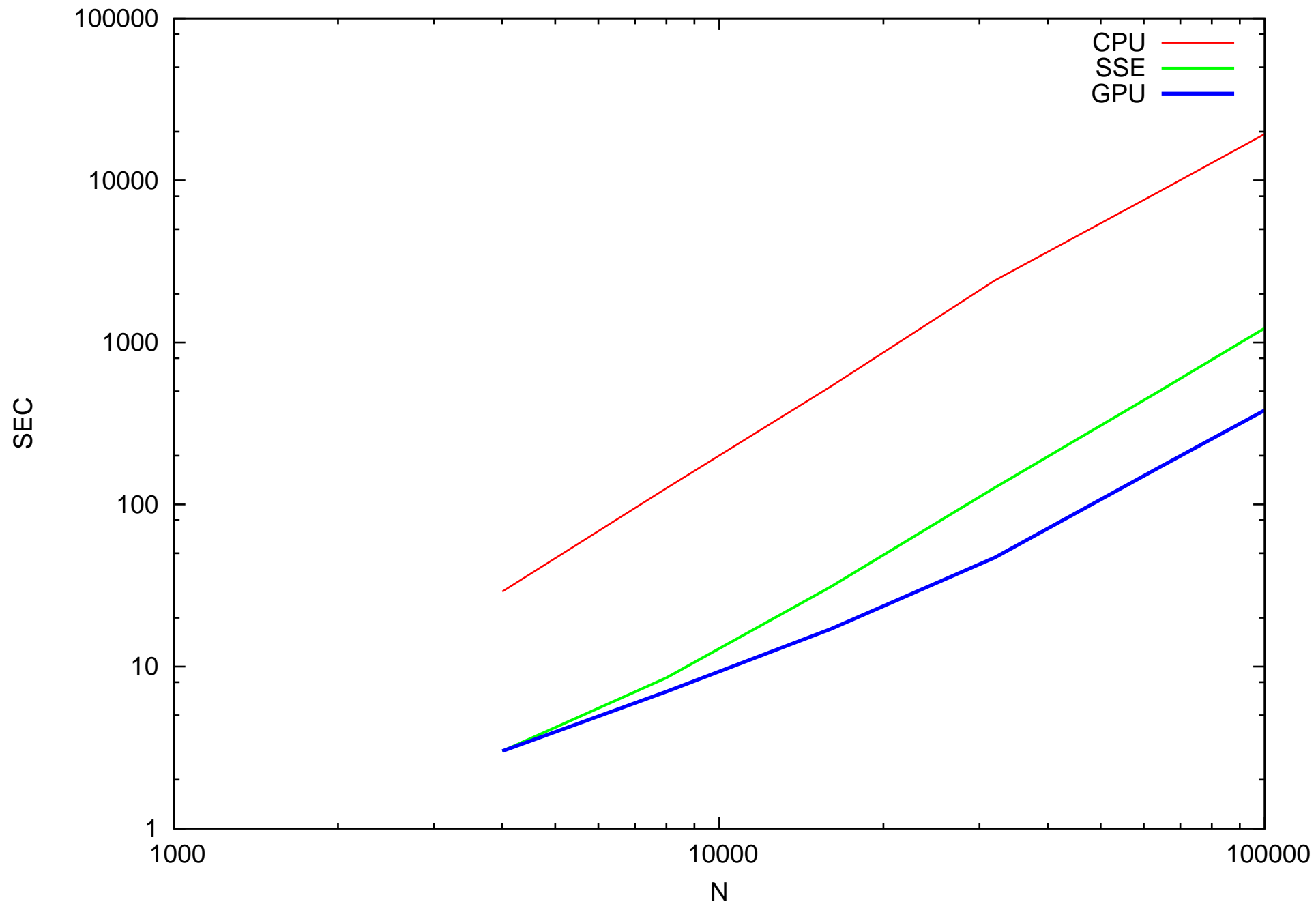
Binding energy $h = \frac{1}{2}\dot{x}^2 - \frac{M}{x}$

Substitution $\dot{x} = \frac{x'}{x} \Rightarrow x'' = 2hx + M$

Coordinate transformation $u^2 = x$

Twice diff. of u^2 and $h \Rightarrow u'' = \frac{1}{2}hu$

Regular equation for $x \Rightarrow 0$



Neighbour Scheme

Total force $\mathbf{F}(t) = \sum_{j=1}^n \mathbf{F}_j + \mathbf{F}_d(t)$

Prediction & $\dot{\mathbf{F}}$

$$\mathbf{F}(t) = \mathbf{F}_n + \dot{\mathbf{F}}_d(t - t_0) + \mathbf{F}_d(t_0), \quad \dot{\mathbf{F}} = \dot{\mathbf{F}}_n + \dot{\mathbf{F}}_d$$

Time-scales $\Delta t_n \ll \Delta t_d, \quad n \ll N$

Contrast $C = \frac{2n}{N} \left(\frac{r_h}{R_s} \right)^3, \quad n_p = n_{max}(0.04 * C)^{1/2}$

Neighbour sphere $R_s^{\text{new}} = R_s^{\text{old}} \left(\frac{n_p}{n} \right)^{1/3}$

Neighbour selection $|\mathbf{r}_i - \mathbf{r}_j| < R_s, \quad \Rightarrow \text{list}$

Derivative corrections $\mathbf{F}_{ij}^{(2)}, \mathbf{F}_{ij}^{(3)}$

Time-Steps

Basic time-step $\Delta t = \frac{\alpha|\mathbf{r}|}{|\mathbf{v}|}, \quad \Delta t = \frac{\beta|\mathbf{F}|}{|\mathbf{F}^{(1)}|}$

Taylor series $\mathbf{F} = \mathbf{F}_0 + \mathbf{F}_0^{(1)} \Delta t + \frac{1}{2}\mathbf{F}_0^{(2)} \Delta t^2 + \dots$

Natural time-step $\Delta t = \left(\frac{\eta|\mathbf{F}|}{|\mathbf{F}^{(2)}|} \right)^{1/2}, \quad \eta = 0.02$

General expression $\Delta t = \left(\frac{\eta(|\mathbf{F}||\mathbf{F}^{(2)}| + |\mathbf{F}^{(1)}|^2)}{|\mathbf{F}^{(1)}||\mathbf{F}^{(3)}| + |\mathbf{F}^{(2)}|^2} \right)^{1/2}$

Relative criterion Δt independent of mass

Block-steps $\Delta t_n = \frac{\Delta t_1}{2^{n-1}}, \quad \Delta t_1 = 1$

Hierarchical levels \mathcal{N}_k particles with steps Δt_k

Scheduling $i = \min (t_j + \Delta t_j)$

Levi-Civita Formulation

2D system: u_1, u_2

$$\begin{aligned} R_1 &= u_1^2 - u_2^2 \\ R_2 &= 2u_1u_2 \end{aligned}$$

Transformation

$$\mathbf{R} = \mathcal{L}(\mathbf{u})\mathbf{u}$$

Levi-Civita [1920] matrix

$$\mathcal{L}(\mathbf{u}) = \begin{bmatrix} u_1 & -u_2 \\ u_2 & u_1 \end{bmatrix} \quad \Rightarrow \quad R = u_1^2 + u_2^2$$

Definition $\dot{\mathbf{R}}^2 = \dot{\mathbf{R}}^T \cdot \dot{\mathbf{R}}$ with $\mathbf{R}' = 2\mathcal{L}(\mathbf{u})\mathbf{u}'$ and $\dot{R} = R'/R$

$$\dot{\mathbf{R}} = 2\mathcal{L}(\mathbf{u})\mathbf{u}'/R$$

$\dot{\mathbf{R}}^T = 2\mathbf{u}'\mathcal{L}^T(\mathbf{u})/R$ and $\mathcal{L}^T(\mathbf{u})\mathcal{L}(\mathbf{u}) = R\mathbf{I}$ give

$$\dot{\mathbf{R}}^T \cdot \dot{\mathbf{R}} = 4\mathbf{u}' \cdot \mathbf{u}'/R$$

From $\mathcal{L}'(\mathbf{u}) = \mathcal{L}(\mathbf{u}')$ we have $\mathbf{R}'' = 2\mathcal{L}(\mathbf{u})\mathbf{u}'' + 2\mathcal{L}(\mathbf{u}')\mathbf{u}'$

Final equation of motion, with $\mathbf{u} \cdot \mathbf{u} = R$

$$\mathbf{u}'' = \frac{1}{2}h\mathbf{u} + \frac{1}{2}R\mathcal{L}^T(\mathbf{u})\mathbf{F}_{kl}$$

Binding energy per unit reduced mass

$$h = [(2\mathbf{u}' \cdot \mathbf{u}' - (m_k + m_l)]/R$$

Rate of change from $\dot{\mathbf{R}} \cdot \ddot{\mathbf{R}}$

$$\frac{d}{dt} \left[\frac{1}{2}\dot{\mathbf{R}}^2 - \frac{(m_k + m_l)}{R} \right] = \dot{\mathbf{R}} \cdot \mathbf{F}_{kl}$$

Conversion by $h' = \mathbf{R}' \cdot \mathbf{F}_{kl}$ and $\dot{\mathbf{R}}$

$$h' = 2\mathbf{u}' \cdot \mathcal{L}^T(\mathbf{u})\mathbf{F}_{kl}$$

KS Regularization

New coordinates $R = u_1^2 + u_2^2 + u_3^2 + u_4^2$

Time transformation $dt = R d\tau$

Coordinate transformation $\mathbf{R} = \mathcal{L}(\mathbf{u}) \mathbf{u}$

Levi-Civita matrix

$$\mathcal{L}(\mathbf{u}) = \begin{bmatrix} u_1 & -u_2 & -u_3 & u_4 \\ u_2 & u_1 & -u_4 & -u_3 \\ u_3 & u_4 & u_1 & u_2 \end{bmatrix}$$

Equations of motion

$$\begin{aligned} \mathbf{u}'' &= \frac{1}{2} h \mathbf{u} + \frac{1}{2} R \mathcal{L}^T \mathbf{P} \\ h' &= 2 \mathbf{u}' \cdot \mathcal{L}^T \mathbf{P} \\ t' &= \mathbf{u} \cdot \mathbf{u} \end{aligned}$$

Close encounter $\Delta t_i < \Delta t_{cl}; \quad R < r_{cl}$

Termination $\gamma \equiv \frac{|\mathbf{P}| R^2}{m_i + m_j} > 0.5$

Centre of mass motion $\dot{\mathbf{r}} = \frac{m_i \mathbf{P}_i + m_j \mathbf{P}_j}{m_i + m_j}$

Perturber selection $r_k < \lambda R, \quad \gamma > 1 \times 10^{-6}$

KS Initialization

Case (i) $R_1 \leq 0$: add R_1 to R ,

$$u_1^2 + u_4^2 = \frac{1}{2}(R_1 + R).$$

Redundancy choice $u_4 = 0$ gives

$$u_1 = [\frac{1}{2}(R_1 + R)]^{1/2},$$

$$u_2 = \frac{1}{2}R_2/u_1,$$

$$u_3 = \frac{1}{2}R_3/u_1.$$

Case (ii) $R_1 < 0$: subtract R_1 from R ,

$$u_2^2 + u_3^2 = \frac{1}{2}(R - R_1).$$

Setting $u_3 = 0$ leads to

$$u_2 = [\frac{1}{2}(R - R_1)]^{1/2},$$

$$u_1 = \frac{1}{2}R_2/u_2,$$

$$u_4 = \frac{1}{2}R_3/u_2.$$

Derivative of \mathbf{R} and property of $\mathcal{L}(\mathbf{u})$ yields

$$\mathbf{R}' = 2\mathcal{L}(\mathbf{u})\mathbf{u}'.$$

Apply $\mathcal{L}^T(\mathbf{u})$ on both sides with $\mathcal{L}^T\mathcal{L} = R$,

$$\mathbf{u}' = \frac{1}{2}\mathcal{L}(\mathbf{u})\mathbf{R}'/R.$$

Definition $R' = R\dot{R}$ gives the KS velocity

$$\mathbf{u}' = \frac{1}{2}\mathcal{L}(\mathbf{u})\dot{\mathbf{R}}.$$

Hermite KS

Standard KS

$$\begin{aligned}\mathbf{u}'' &= \frac{1}{2}h \mathbf{u} + \frac{1}{2}R \mathcal{L}^T \mathbf{F}_{kl} \\ h' &= 2 \mathbf{u}' \cdot \mathcal{L}^T \mathbf{F}_{kl} \\ t' &= \mathbf{u} \cdot \mathbf{u}\end{aligned}$$

New notation

$$\begin{aligned}\mathbf{F}_u &= \mathbf{u}'' \\ \mathbf{Q} &= \mathcal{L}^T \mathbf{P},\end{aligned}$$

with $\mathbf{P} = \mathbf{F}_{kl}$ as the perturbing force.

Basic equations

$$\begin{aligned}\mathbf{F}_u &= \frac{1}{2}h \mathbf{u} + \frac{1}{2}R \mathbf{Q} \\ h' &= 2 \mathbf{u}' \cdot \mathbf{Q} \\ t' &= \mathbf{u} \cdot \mathbf{u}\end{aligned}$$

Hermite \mathbf{F} , \mathbf{F}' formulation

$$\begin{aligned}\mathbf{F}_u &= \frac{1}{2}h \mathbf{u} + \frac{1}{2}R \mathbf{Q} \\ \mathbf{F}'_u &= \frac{1}{2}(h' \mathbf{u} + h \mathbf{u}' + R' \mathbf{Q} + R \mathbf{Q}') \\ h' &= 2 \mathbf{u}' \cdot \mathbf{Q} \\ h'' &= 2 \mathbf{F}_u \cdot \mathbf{Q} + 2 \mathbf{u}' \cdot \mathbf{Q}' \\ t' &= \mathbf{u} \cdot \mathbf{u}\end{aligned}$$

The derivatives of \mathbf{P} , \mathbf{Q} and t' are readily available. Note that $\mathbf{P}' = R \dot{\mathbf{P}}$ and that $\mathcal{L}^T(\mathbf{u}')$ can be obtained by substituting \mathbf{u}' for \mathbf{u} . For implementation, significant accuracy can be gained by high-order prediction (not used in standard Hermite).

KS Decision-Making

Close encounter	$R_{\text{cl}} = \frac{4 r_h}{N C^{1/3}}, \quad \Delta t_{\text{cl}} = \beta \left(\frac{R_{\text{cl}}^3}{\bar{m}} \right)^{1/2}$
Time-step criterion	$\Delta t_k < \Delta t_{\text{cl}}$
Neighbour list search	list all $r_{kj}^2, \quad \Delta t_j < 2 \Delta t_{\text{cl}}$
Two-body selection	$R_{kl} < R_{\text{cl}}, \quad \dot{R}_{kl} < 0$
Dominant motion	$\frac{m_k + m_l}{R_{kl}^2} > \frac{m_k + m_j}{R_{kj}^2}$
KS initialization	$\mathbf{F}_U, \mathbf{F}'_U, \Delta\tau \ \& \ t^{(n)} \Rightarrow \Delta t$
Initialization of c.m.	$\mathbf{r}_{\text{cm}} = \frac{m_k \mathbf{r}_k + m_l \mathbf{r}_l}{m_k + m_l}$
Perturber search	$r_p < \left(\frac{2m_p}{m_b \gamma_{\text{min}}} \right)^{1/3} a (1 + e)$
Slow-down adjustment	$\gamma < \gamma_0, \quad \Delta\tau \Rightarrow \kappa \Delta t$
Termination test	$R > R_0, \quad \gamma > \gamma^*$
Delayed termination	$T_{\text{block}} - t > \Delta t_i$
Final iteration	$\Delta\tau$ from $\dot{\tau}, \ddot{\tau}, \dots$ and δt
Polynomial initialization	$\mathbf{F}_j, \dot{\mathbf{F}}_j, \Delta t_j, \quad j = k, l$

Practical Aspects of KS

Regular equations	Perturbed harmonic oscillator, $\gamma < 1$
Constant time-step	$\Delta\tau = \eta \left(\frac{1}{2 h } \right)^{1/2}$ vs $\Delta t \propto R^{3/2}$
Linearized equations	Higher accuracy per step
Faster force calculation	Tidal perturbation, $P \propto 1/r^3$
Unperturbed motion	$\gamma < 10^{-6}$, $\Delta t > t_K$
Slow-down procedure	Adiabatic invariance, $\tilde{P} = \kappa P$
Energy rectification	Improve \mathbf{u} , \mathbf{u}' from integration of h'
C.m. approximation	$d > 100 a (1 + e)$
Transformations	$\mathbf{R} = \mathcal{L}\mathbf{u}$, $\mathbf{r}_j = \mathbf{r}_{\text{cm}} \pm \mu\mathbf{R}/m_j$ $\dot{\mathbf{R}} = 2\mathcal{L}\mathbf{u}'/R$, $\dot{\mathbf{r}}_j = \dot{\mathbf{r}}_{\text{cm}} \pm \mu\dot{\mathbf{R}}/m_j$
Two-body elements	a, \mathbf{J}, e for averaging & circularization

Scaling of Initial Conditions

Main input	$N, N_b, M_S, R_{\text{pc}}$
Cluster parameters	optional IMF and Plummer or King model
Initial data	$\tilde{m}_i, \tilde{\mathbf{r}}_i, \tilde{\mathbf{v}}_i, \dots, i = 1, N$
Total energy	$E = T - U$
Virial theorem	$\mathbf{v}_i = q \tilde{\mathbf{v}}_i, \quad q = \left[\frac{Q_V U}{T} \right]^{1/2}, \quad \mathbf{r}_i = \tilde{\mathbf{r}}_i$
Standard units	$G = 1, \quad \Sigma m_i = 1, \quad E_0 = -0.25$
Standard scaling	$\hat{\mathbf{r}}_i = \frac{\mathbf{r}_i}{S^{1/2}}, \quad \hat{\mathbf{v}}_i = \mathbf{v}_i S^{1/2}, \quad S = \frac{E_0}{q^2 T - U}$
Astrophysical units	V^*, T^*, R^* from $M_{\text{tot}}, R_{\text{pc}}$
Primordial binaries	split or copy m_i , introduce a, e, Ω
Force polynomials	$\mathbf{F}_i, \dot{\mathbf{F}}_i, \Delta t_i, \dots, i = 1, N$
KS regularization	explicit initialization, $R < R_{\text{cl}}$

Units

(a) Scaling relations

Given length scale R_V in pc and total mass NM_S in M_\odot

Velocity scaling

$$\tilde{V}^* = 1 \times 10^{-5} (GM_\odot/L^*)^{1/2} \text{ km/s, with } L^* = 3 \times 10^{18} \text{ cm}$$

$$\text{Velocity unit} \quad V^* = 6.557 \times 10^{-2} (NM_S/R_V)^{1/2} \text{ km/s}$$

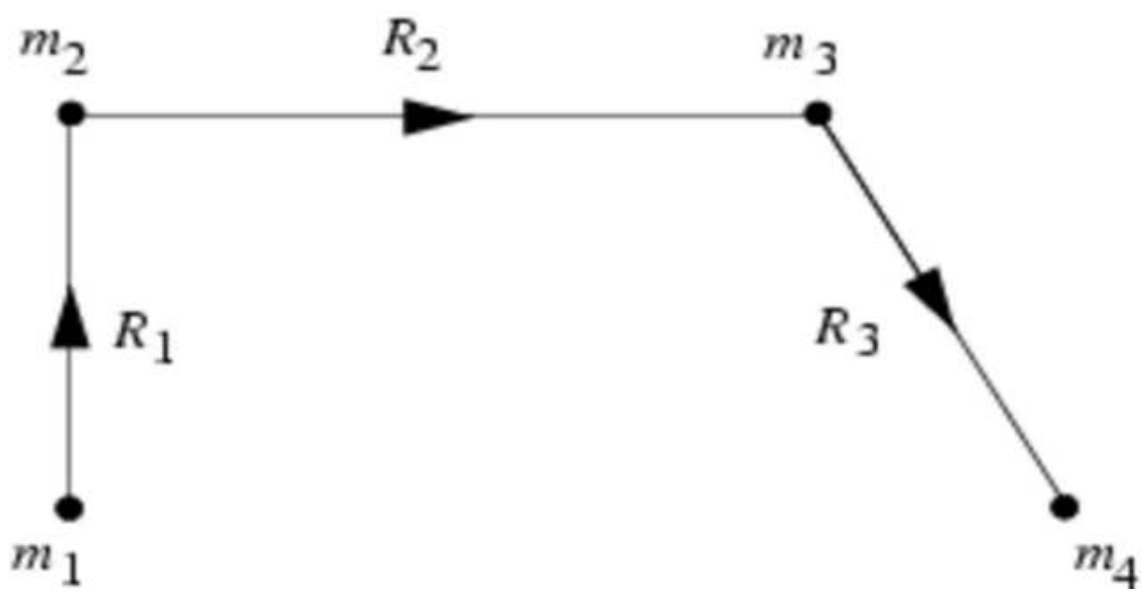
$$\text{Fiducial time} \quad \tilde{T}^* = (L^{*3}/GM_\odot)^{1/2} = 14.94 \text{ Myr}$$

$$\text{Time unit} \quad T^* = 14.94 (R_V^3/NM_S)^{1/2} \text{ Myr}$$

(b) Conversion from N-body to physical units

$$\tilde{r} = R_V r \text{ pc, } \tilde{v} = V^* v \text{ km/s, } \tilde{t} = T^* t \text{ Myr,}$$
$$\tilde{m} = NM_S m M_\odot$$

$$\text{Crossing time} \quad T_{\text{cr}} = 2\sqrt{2} T^* \text{ Myr}$$



Data Structure

Single stars	$2N_p < i \leq N, \quad \mathcal{N}_i = i$
KS pairs	$1 \leq i \leq 2N_p, \quad i_p = i_{\text{icm}} - N$
C.m. particles	$i > N, \quad \mathcal{N} = N_0 + \mathcal{N}_k, \quad k = 2i_p - 1$
Stable triples	KS + ghost, $\mathcal{N}_{\text{cm}} = -\mathcal{N}_k$
Ghost particles	$\mathcal{N}_{\text{ghost}} = \mathcal{N}_{2i_p-1}, \quad m_{\text{ghost}} = 0$
Stable quadruples	KS + KS ghost, $\mathcal{N}_{\text{cm}} = -\mathcal{N}_k$
Higher orders	T + KS, $\mathcal{N}_{\text{cm}} = -(2N_0 + \mathcal{N}_k)$
Chain members	$2N_p < i_{\text{cm}} \leq N, \quad \mathcal{N}_{\text{cm}} = 0$
Single escape	$2N_p < i \leq N, \quad r_i > 2r_{\text{tide}}, \quad \text{remove } i$
Binary escape	$i > N, \quad r_i > 2r_{\text{tide}}, \quad 2i_p - 1, 2i_p$
Hierarchy escape	$i > N, \quad r_i > 2r_{\text{tide}}, \quad 2i_p - 1, 2i_p, i_{\text{ghost}}$

Energy Budget

Definition of total energy

$$E_{\text{tot}} = T + U + E_{\text{tide}} + E_{\text{bin}} + E_{\text{merge}} + E_{\text{coll}} + E_{\text{mdot}} + E_{\text{cdot}} + E_{\text{ch}} + E_{\text{sub}}$$

T Kinetic energy of single bodies and c.m. particles

U Potential energy of single and c.m. bodies

E_{tide} Tidal energy due to external perturbations

E_{bin} Binding energy in regularized pairs

E_{merge} Total internal energy of hierarchical systems

E_{coll} Sum of binding energies released in collisions

E_{mdot} Energy change from mass loss and Roche mass transfer

E_{cdot} Neutron star kicks and common envelope evolution

E_{ch} Total energy of any existing chain subsystem

E_{sub} Energy of unperturbed triple and quadruple subsystems

ΔE Energy change due to removal of escapers

NBODY6 Input File

1 20.0

1000 1 5 50000 95 1

0.01 0.01 0.3 2.0 10.0 100.0 1.0D-04 1.0 0.5

0 0 0 0 1 0 1 0 0 0

0 0 0 1 1 1 0 1 3 4

1 0 2 0 0 0 0 0 0 1

0 0 0 0 0 0 0 0 0 1

0 0 0 0 0 0 0 0 0 0

4.0D-05 4.0D-04 0.2 1.0 1.0D-06 0.001

2.3 10.0 0.2 0 0 0.02 0.0 100.0

0.5 0 0 0 0.125

KSTART TCOMP

N NFIX NCRIT NRAND NNBMAX NRUN

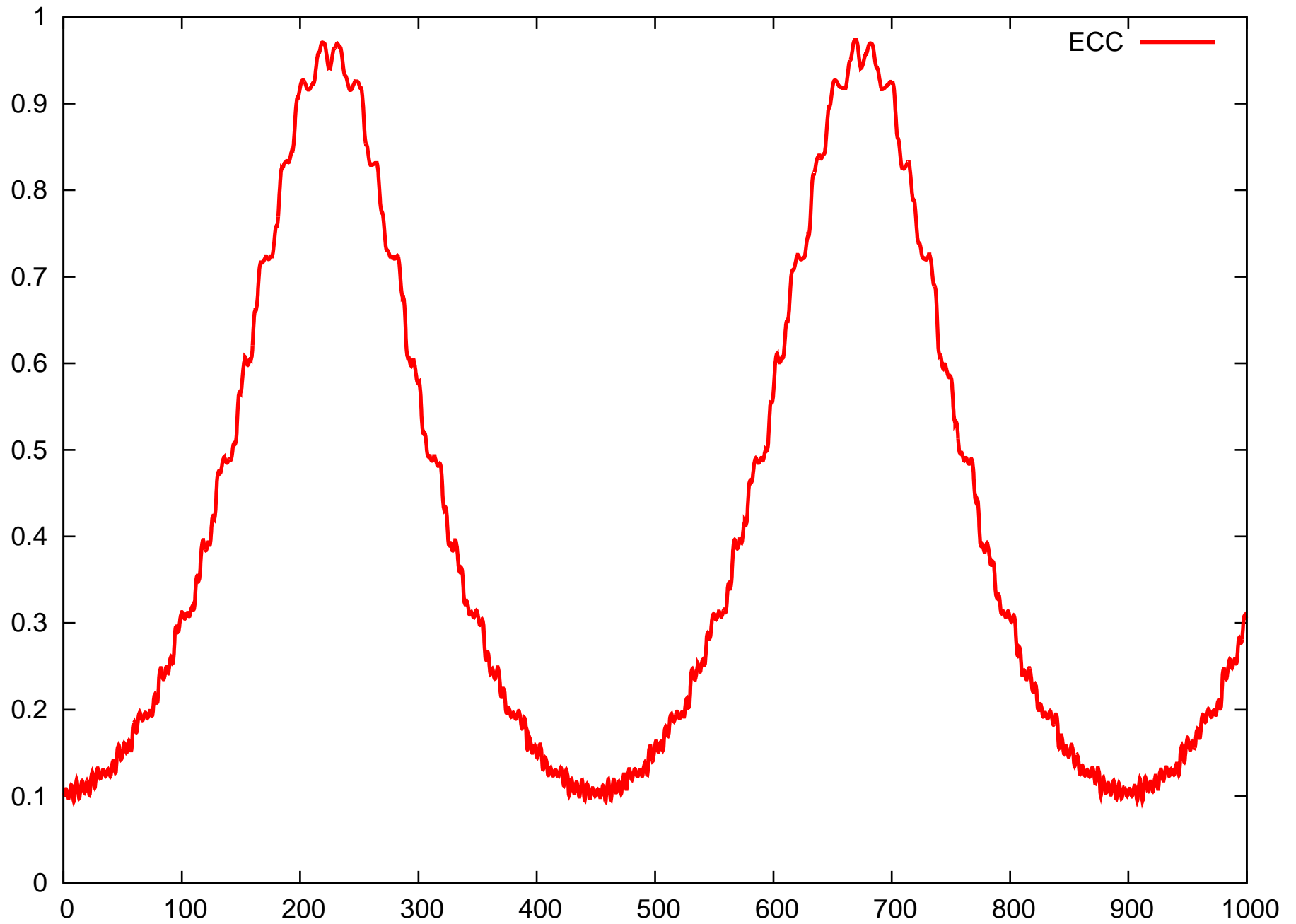
ETAI ETAR RS0 DTADJ DELTAT TCRIT QE RBAR ZMBAR

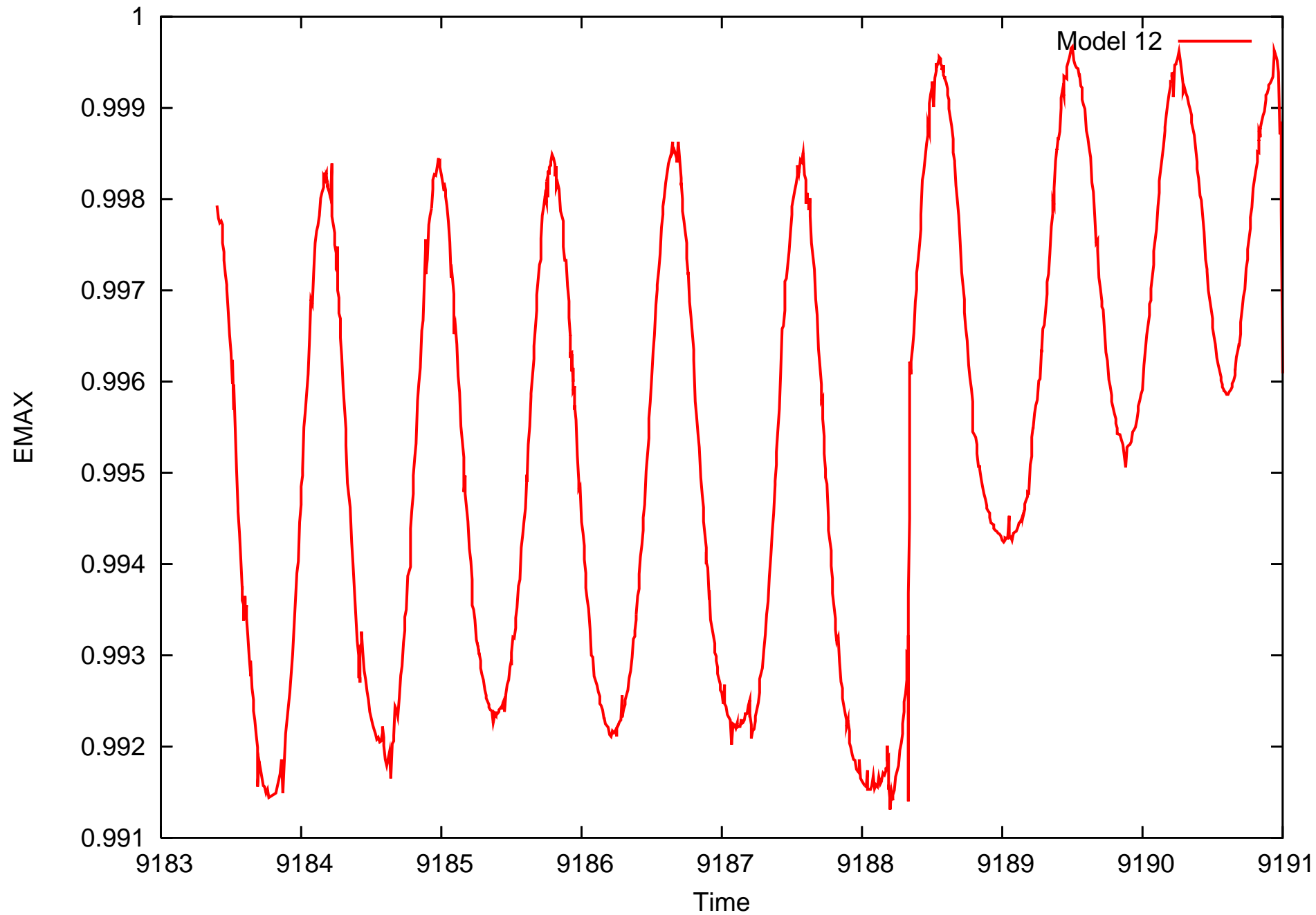
OPTIONS (50)

DTMIN RMIN ETAU ECLOSE GMIN GMAX

ALPHA BODY1 BODYN NBIN0 NHI0 ZMET EPOCH0 DTPLOT

Q 0 0 0 SMAX





Three-Body Stability Criterion

```
REAL*8 FUNCTION QSTAB(e,eout, zi, m1, m2, m3)
IMPLICIT REAL*8 (A-H,O-Z)
REAL*8 m1,m2,m3
*
* Three-body stability function (Valtonen, AAS 2015).
*
* Inner and outer eccentricity: e & eout.
* Inclination in radians: zi.
* Masses: inner, m1 m2, outer, m3.
*
* Adopt 10,000 outer orbits for random walk time-scale.
ZN = 10000.0
zz = 1.0/6.0D0
*
F = 1.0 - 2.0*e/3.0*(1.0 - 0.5*e**2) - 0.3*cos(zi)*
& (1.0 - 0.5*e + 2.0*cos(zi)*(1.0 - 2.5*e**1.5 - cos(zi)))
*
G = SQRT(m1/(m1 + m2))*(1.0 + m3/(m1 + m2))
*
QSTAB = 1.52*(SQRT(ZN)/(1.0 - eout)) **zz * (F * G) **0.3333

RETURN
END
```

Hierarchical Stability

Requirement	a_0 secularly constant
Kozai cycles	$e_{\max} = \left(1 - 5\cos^2 i/3\right)^{1/2}$
Candidates	$\Delta t_{\text{cm}} < \Delta t_{\text{cl}}, a_1(1 - e_1) > 3a_0$
Restrictions	$\mu_1 M_{123}/2a_1 > E_{\text{hard}}, \gamma_1 < 0.01$
Stability test	$f(a_0, a_1, e_0, e_1, \phi, m_1, m_2, m_3)$
Data structure	New KS, m_3 + inner c.m.
Merger table	$m_1, m_2, \mathbf{R}, \mathbf{V}, h, \mathbf{u}, \mathbf{u}', \mathcal{N}_g, \mathcal{N}_{\text{cm}}$
Initialization	New polynomials for KS and c.m.
Assessment	New check $R_{\text{apo}} < P_{\text{crit}}$
Mass loss	Update $h, \mathbf{u}, \mathbf{u}'$, stability check
Termination	$\gamma > 0.1, R > R_{\text{cl}}$ or $\gamma > 0.25$
Re-initialize	Triple \Rightarrow KS + m_3

Physical Collisions

Simple definition	$R_{\text{coll}} = \frac{3}{4}(r_1^* + r_2^*)$
Two-body encounter	KS regularization
Pericentre condition	$R'_0 R' < 0, \quad R < a$
Pericentre determination	Δt_{peri} from Kepler's equation
Predict \mathbf{R}_{peri} or iterate	$d\tau_0 = \frac{\Delta t_{\text{peri}}}{R}$, Newton-Raphson
Implement collision	$m_{\text{cm}} = m_1 + m_2, \quad r_{\text{cm}} = \frac{m_1 \mathbf{r}_1 + m_2 \mathbf{r}_2}{m_1 + m_2}$
Mass loss	$\Delta m = f(K_1^*, K_2^*)$
Initialize single body	$\mathbf{F}_1, \dot{\mathbf{F}}_1, \Delta t_1$
Compact subsystem	$\dot{R} \simeq 0$ by iteration
Transformation	$\mathbf{Q}, \mathbf{P} \Rightarrow \mathbf{r}, \dot{\mathbf{r}}$
New chain construction	$N_{\text{ch}} \Rightarrow N_{\text{ch}} - 1, \quad E_{\text{coll}} = E_{\text{ch}} - \mathcal{V}$

Simulations

Initial conditions $N = 1 \times 10^5$, Kroupa IMF in $0.1 - 50M_{\odot}$

Plummer & tides $R_h = 2$ pc, $V^* = 10$ km/s, $T^* = 0.2$ Myr

N-body units $G = 1$, $\bar{m} = 1/N$, $E = -\frac{1}{4}$, $\bar{v}^2 = \frac{1}{2}$

Hard binary $\frac{m^2}{2a} \simeq \frac{1}{2}\bar{m}\bar{v}^2$, $\Rightarrow a_{\text{hard}} \simeq 2/N$

Super-hard $-\frac{m^2}{2a} = E$, $\Rightarrow a_0 = 2/N^2$, or 2×10^{-10}

Dynamics $m_{\text{bh}} \simeq 20 M_{\odot}$, $E_b = 0.01E$, $a \simeq 3 \times 10^{-5}$

Coalescence $R_{\text{co}} = \frac{8(m_1 + m_2)}{c^2}$, $R_{\text{co}} \simeq 7 \times 10^{-12}$

PN condition $a(1 - e) \simeq 10^4 R_{\text{co}}$, $\Rightarrow e > 0.999$

Time-scale $a = 3 \times 10^{-5}$, $e = 0.999$, $m = 20 M_{\odot}$, $\tau_{\text{co}} \simeq 200$ Myr

Einstein shift $\Delta w = \frac{6\pi M}{ac^2(1 - e^2)} = 3 \times 10^{-4}$

Getting Started

1. Download code `nbody6.tar.gz`
2. Unzip `gunzip nbody6.tar.gz`
3. Extract files `tar xvf nbody6.tar`
4. Check `params.h` `NMAX, LMAX, KMAX, MMAX`
5. Compile the code `make nbody6`
6. Create run directory `mkdir Run`
7. Run test input `time nbody6 <input >output &`
8. Profiling `Makefile with -O3 -pg`
9. Performance data `gprof nbody6 gmon.out -p >OUT`