Meridional flux transport on the Sun

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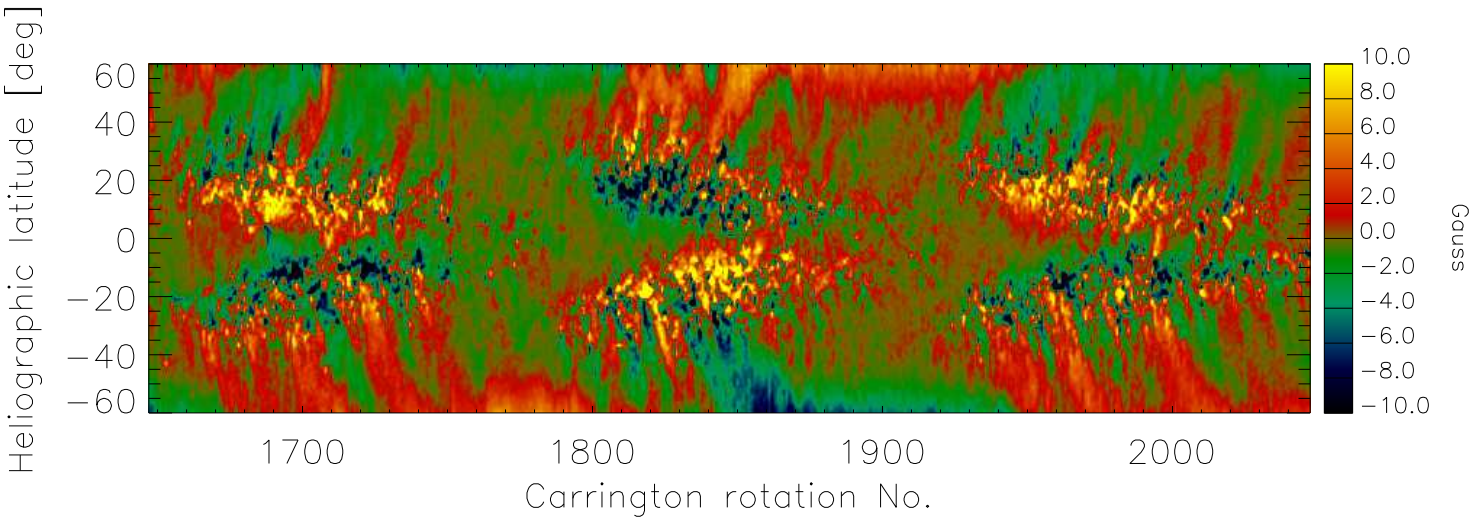
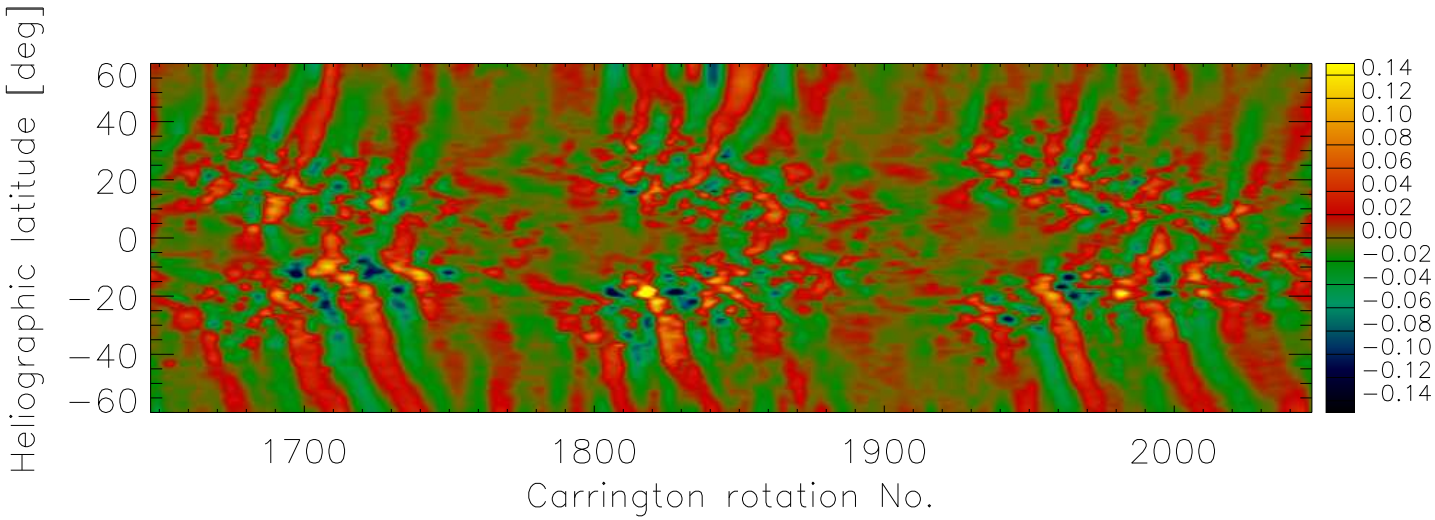
Introduction

- Flux transport essential for dynamo models
- Wang & Sheeley's works
  - Stream of magnetic flux spread poleward from the sunspot belts, deflected eastward by differential rotation.
  - Two components: supergranular diffusion and meridional transport
  - Without MT – no polar fields

Flux transport in latitude and time



Frequency filter 1.06×10^{-8}–3.17×10^{-7} s^{-1} (periods 36.5–1093 days)



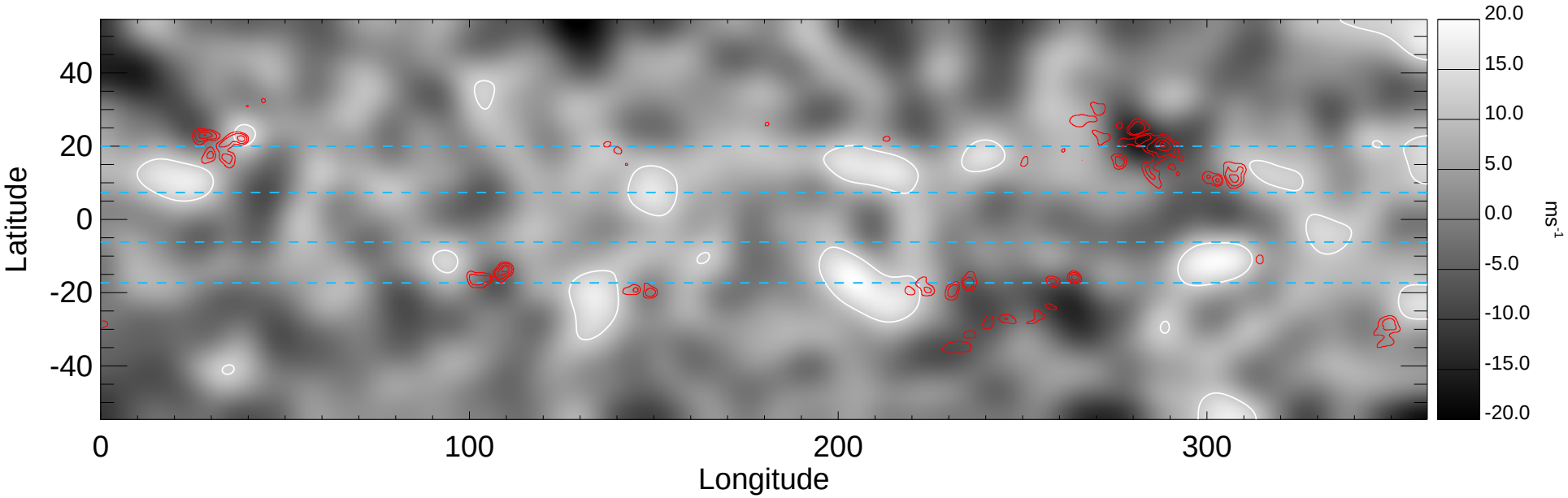
Still in progress. . .

Some questions:

- Is this important for the sub-surface layer dynamics?
- Could this contribute to the formation of torsional oscillations?
- How the torsional oscillations look like? Can we see them in flow maps?

- We are working on it! Sometimes, somehow, . . .

Torsional oscillations, where are you?



Testovaci

Kulaty box

Toto je text v kuletem boxu se stinem.

Druhy Kulaty box

Toto je text v druhem kuletem boxu se stinem.

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Odkryvani

- Polozka 1
- Polozka 2
- Polozka 3

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- Polozka 1
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- Polozka 1
- Polozka 2
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Jine odkryvani

- Polozka 1
- Polozka 2
- Polozka 3

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Jine odkryvani

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- Polozka 2
- Polozka 3

Zvyraznovani

- Prvni polozka
- Druha polozka
- Treti polozka
- Ctvrta polozka

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Urceni

- Prvni polozka na prvnim slajdu
- Druha polozka
- Treti polozka
- Ctvrta polozka

Urceni

- Prvni polozka na druhem slajdu
- Druha polozka
- Treti polozka
- Ctvrta polozka

Urceni

- Prvni polozka na tretim slajdu
- Druha polozka
- Treti polozka
- Ctvrta polozka

Urceni

- Prvni polozka na ctvrtem slajdu
- Druha polozka
- Treti polozka
- Ctvrta polozka